



- User's Guide -  
(for version 1904)

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Citation of [O. Vallcorba & J. Rius. d2Dplot: 2D X-ray diffraction data processing and analysis for through-the-substrate microdiffraction J. Appl. Cryst. 2019, 52, 478-484](#) would be greatly appreciated when this program helped to your work.

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## 1. Installation and use of *d2Dplot*

No installation of the program is required. Only extract the files and folders of the zip file into the desired folder in your hard drive and run the executable file (`d2Dplot.exe` in Windows and `d2Dplot` in Linux). In most of the recent Linux distributions, the executable files can be executed by double click from the file explorer but alternatively you can also run it from the command line with `./d2Dplot`. If the execute flag of the file is turned off, turn it on with: `chmod +x d2Dplot`

**Tip:** Running it from the command line has the advantage that you can give an image file as the argument and it will be automatically opened. Also you can use the *macro* mode to give instructions to operate the program through command line arguments (no GUI).

*Note:* JAVA is required (version 1.6.0\_18 or higher).

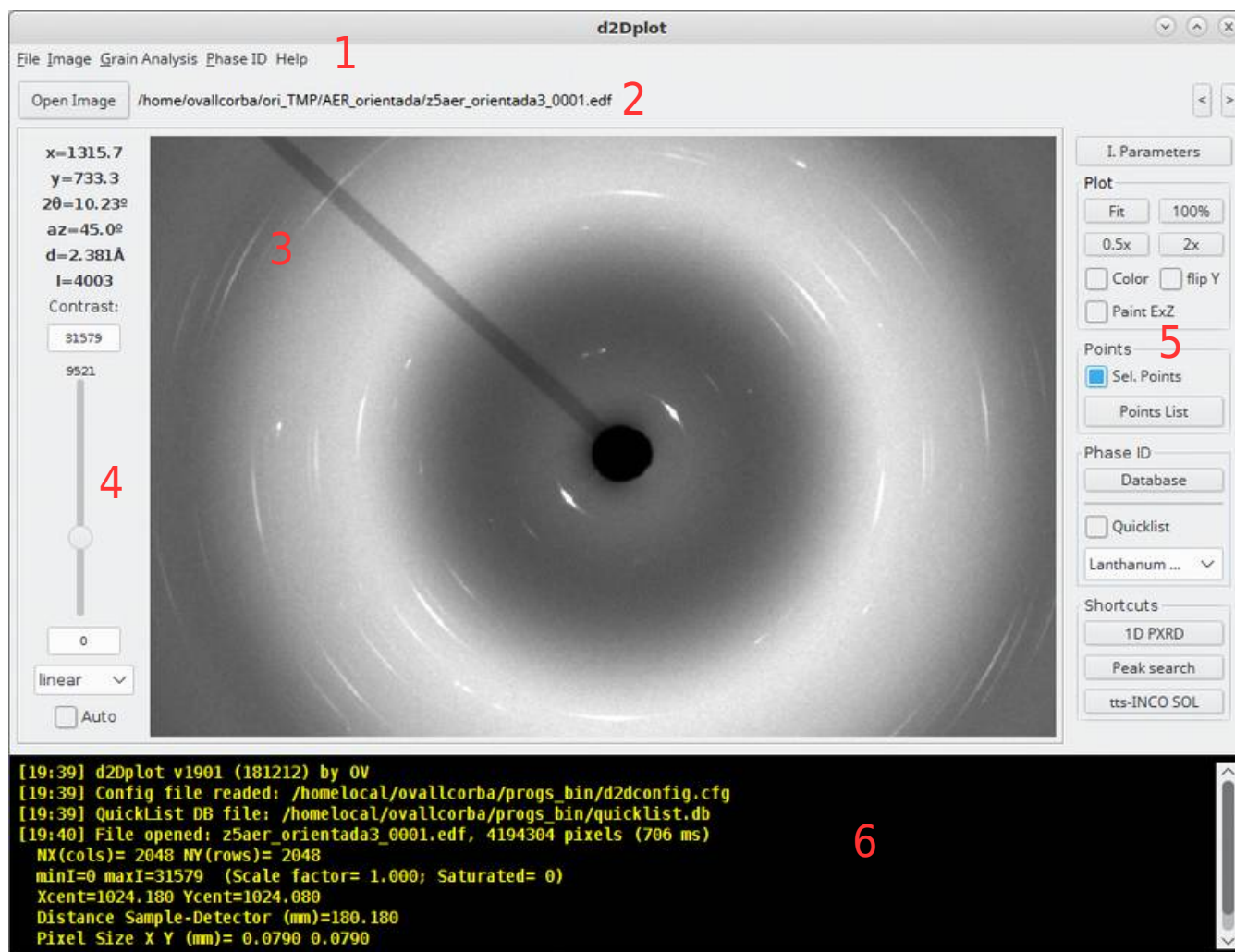
### Configuration file

The first run, the program generates a plain text configuration file (`d2dconfig.cfg`) at the same folder where the program is installed. However, in some systems it can be created inside the user folder or somewhere else (the program will display the location of the file on the output panel located at the bottom part of the main window). Usually there is no need to change anything of this file but, if desired, the parameters are self-explanatory and their value can be modified.

The most important parameters are the default paths to the compound databases (`defQuickListDB` and `defCompoundDB`) which can be modified according to our preferences.

## 2. Overview

This is the aspect of the main window after opening an image (via menu File-Open) or clicking the button Open Image.



The main parts are:

1. **Menu bar.** To access all the program modules and options. It contains:

- **File**
  - Open Image. Opens an image file.
  - Save Image. Save the image file (to any of the supported formats)
  - Export as PNG. Save as a PNG file.
  - Sum Images. To merge several images to a single one.

- **Subtract Images.** To subtract one image from another.
- **Batch Convert.** To convert a list of files to another file format.
- **Fast Viewer.** To open a series of images to be displayed sequentially.
- **Reset.** Resets the program
- **Quit.** Exit the program
- **Image**
  - **Instrumental Parameters.** To introduce the instrumental parameters of the image.
  - **Instr. Param. Calibration.** Use of a standard substance diffraction data (LaB<sub>6</sub>, Silicon, etc...) to calibrate the sample-to-detector distance, the beam center and the tilt/rot of the detector.
  - **Excluded Zones.** To select zones of the image that have to be omitted in further calculations.
  - **Background Subtraction.** To subtract the background of the image. In the case there is some contribution of a holder (glass, etc...) and we want to get a background clean image.
  - **Conversion to 1D PXRD.** To get the corresponding 1D (powder) pattern of the diffraction image.
  - **Azhimuthal (circular) plot.** To get the 1D plot of the intensity along a Debye ring.
  - **HP Cu Pcalc.** To calculate the system pressure from two peaks of the Cu used as internal standard in high pressure experiments.
- **Grain Analysis**
  - **Find/Integrate Peaks.** Locate diffraction peaks on the image.
  - **Run tts\_Software.** Opens the front-end to launch the *TTS\_software*.
  - **Load tts-INCO SOL/PCS files.** Open the files generated by tts-INCO or tts-REDUC programs to check the correctness of single grain orientations (Rius *et. al.* 2015, 2016).
  - **Load XDS file.** Open a spot.xds file from XDS (**X**-ray **D**etector **S**oftware, CCP4; Kabsch, 1988) to show the position of the peaks.
  - **SC data to INCO.** Convert single-crystal dataset (small angular step) to a wider step angle format for tts-INCO.

- **Phase ID**

- **Database.** Opens the compound database window. To plot theoretical rings from a compound database and search compounds from the image ring positions (more explained in the corresponding section of the guide).

- **Help**

- **About.** Some information about the program.
- **Manual.** Link to this user's guide.
- **Check for updates.** To see if a new version of *d2Dplot* is available.

2. **Top bar.** It contains a button to quickly open an image file and also shows the path of the current displayed image. On the right part there are two arrow buttons which allow a quick navigation between consecutive images. Consecutive images are those which have the same filename followed by four sequential digits (e.g. lab6\_0000.d2d, lab6\_0001.d2d, lab6\_0002.d2d, lab6\_0003.d2d,...).

3. **Image panel.** Where the image is shown. The general interaction is:

- Left mouse button: Selection, peak addition, etc... (depend on the opened module)
- Middle mouse button: Press and drag to move the image. Click with no movement to fit the image to the display area.
- Mouse wheel: Zoom.
- Right mouse button: Deletion. Press and drag (UP and DOWN) for zoom.

4. **Image panel controls.** Information about the current pixel we are pointing is shown here. Also the contrast can be adjusted with the slide. The **auto** checkbox is to calculate automatically the contrast value for every opened image (it is done by default on the first one opened but if a consecutive images are opened it is often desired to have it disabled for comparison).

5. **Right panel.** Here, we have:

- Shortcuts for the instrumental parameters, radial integration, grain analysis, peak search/integrate and compound database.
- Plotting options regarding the image display (100% means that a pixel of the screen

corresponds to a pixel of the image). ExZ = Excluded zones.

- Point selection tool. To select (by left mouse button clicking) points or rings of the image. The point list can be retrieved with the button Point List (and exported to a file if it wants to be used somewhere else). These points can be used later to search in the database (or calculate the pressure with Cu).
- Quicklist. It contains a selection of the compounds for quick display of the rings. More about the quicklist is explained in the corresponding section of the guide.

6. **Output panel.** Some messages of the program are displayed here.



### 3. Image menu modules

#### Instrumental parameters

**Instrumental Parameters**

Sample-Detector distance (mm)= 180.18

Pixel size X (mm)= 0.079

Pixel size Y (mm)= 0.079

Beam centre X (pixel)= 1024.18

Beam centre Y (pixel)= 1024.08

Wavelength (Å)= 0.4246

Detector Tilt (°)= 0.0

Detector Rot (°)= 0.0

Saturation value (counts)= 65535

(scan) omega ini (°)= -7.5

(scan) omega end (°)= 7.5

Acquisition time (s)= 30.0

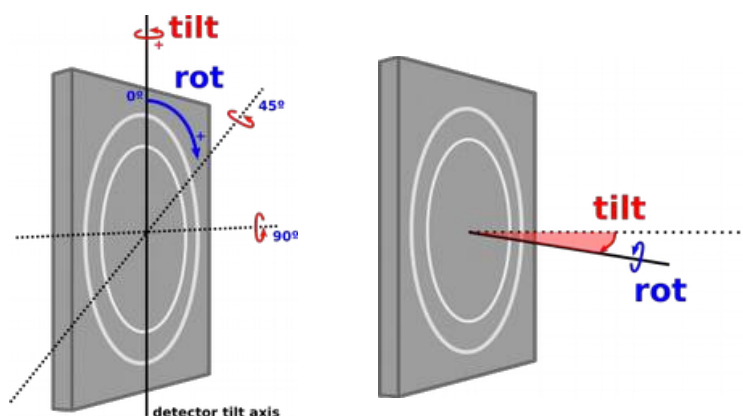
☐ keep calibration info for the session

Update from header Apply Apply and Close Cancel

Instrumental and acquisition parameters are introduced here, names are self explanatory and the units are shown. The Tilt/Rot convention used is:

- Tilt: Deviation (angle) of the orthogonality of the beam direction.
- Rot: Clockwise rotation (angle) of a perpendicular axis taking as “zero” the vertical (i.e. 12h on a clock).

(Better a drawing...)

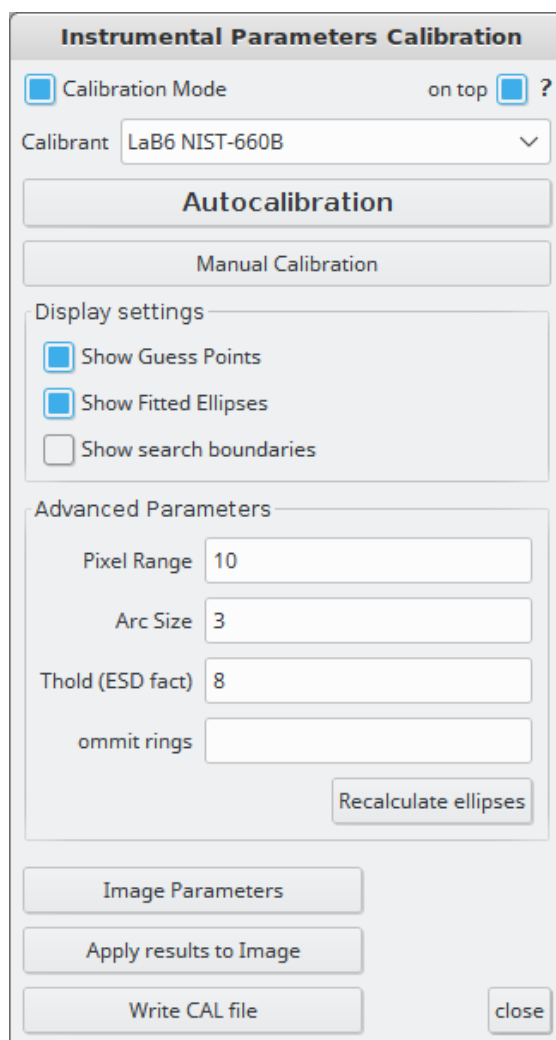


The convention used is compatible with calibrations with *Fit2D* program (Hammersley, Svensson & Thompson, 1994).

In the “?” dialog, a simulated LaB<sub>6</sub> diffraction image can be generated entering custom calibration values.

If *keep calibration info for the session* is selected, no calibration info will be read from the header of next opened images.

## Instrumental Parameters Calibration



The screenshot shows the 'Instrumental Parameters Calibration' dialog box. It has a title bar with the same text. Inside, there's a 'Calibration Mode' section with a checked checkbox and the text 'on top' followed by a question mark icon. Below this is a 'Calibrant' dropdown menu currently set to 'LaB6 NIST-660B'. There are two main buttons: 'Autocalibration' and 'Manual Calibration'. Under 'Autocalibration', there's a 'Display settings' section with three checkboxes: 'Show Guess Points' (checked), 'Show Fitted Ellipses' (checked), and 'Show search boundaries' (unchecked). Below that is an 'Advanced Parameters' section with four input fields: 'Pixel Range' (10), 'Arc Size' (3), 'Thold (ESD fact)' (8), and 'ommit rings' (empty). A 'Recalculate ellipses' button is located to the right of the 'ommit rings' field. At the bottom, there are three buttons: 'Image Parameters', 'Apply results to Image', and 'Write CAL file', along with a 'close' button in the bottom right corner.

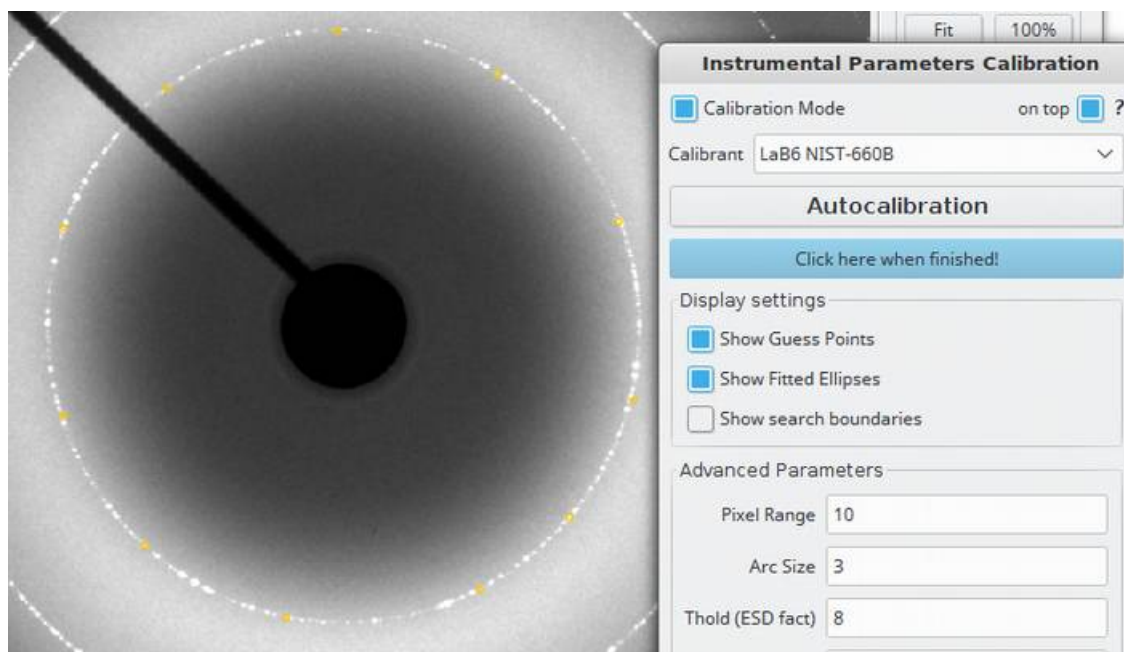
This module is for calibration of the sample-to-detector distance, the beam center and the detector tilt/rot angles from a calibrant substance (e.g. LaB<sub>6</sub>). The calibrant substance can be selected in the same window and there are two by default in *d2Dplot*, LaB<sub>6</sub> and Silicon. However, more calibrants can be added to the list by introducing additional lines in the

config file (d2dconfig.cfg) starting with "calibrant = " followed by an identifying name and a list of  $d$ -spacings separated by semicolons (;), for example:

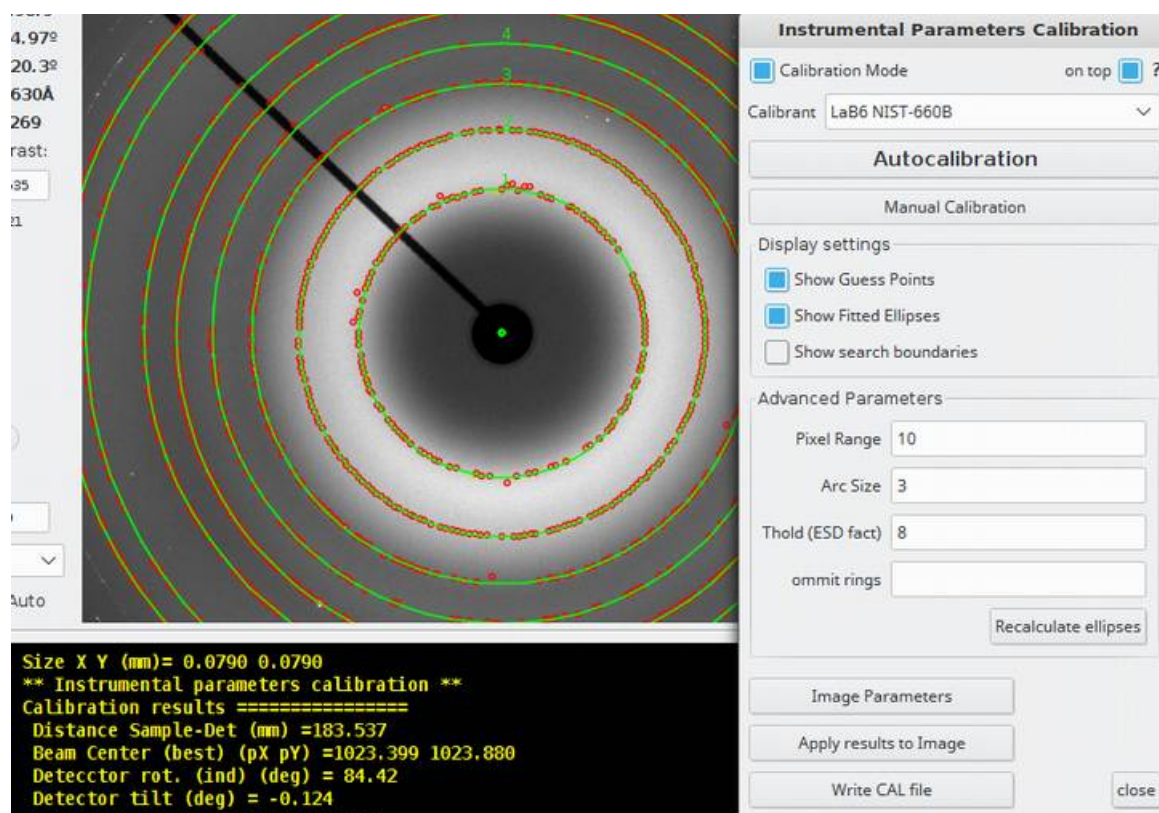
```
calibrant = LaB6 NIST-660B; 4.1568785; 2.9393575; 2.3999755; 2.0784323; 1.8590043
```

To perform the calibration there are two options:

- a) Click on **Autocalibration**. It will use the image header info (distance, wavelength, center, ...) as initial guess values to find the calibrant rings automatically. Try this method first.
- b) Select manually the points on the first ring of the calibrant:
  1. Click on manual calibration button
  2. Click >5 points on the inner calibrant ring (do not need to be very accurate...)



3. Click on the same button (now labeled **Click here when finished!**)
4. The rings and instrumental parameter values will be calculated



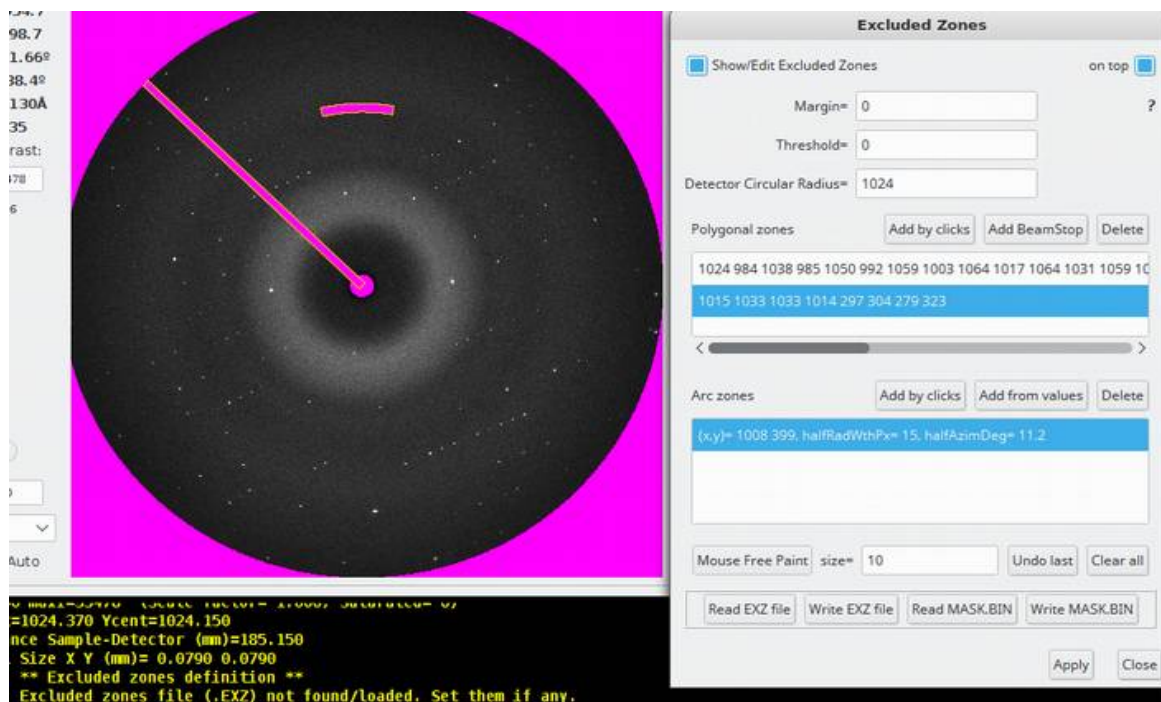
5. It can be repeated by clicking on the **Recalc** button (changing the parameters from the search rings if necessary).

The display settings show more info regarding the search of the calibrant rings.

Buttons below allow to apply the calibration results to the current image or write a CAL file to be used as calibration info for batch processing of images (and/or when the header info is not enough or correct).

The instrumental parameters estimation (beam center, distance, detector tilt and rot) are obtained following the methodology described by Hart *et al.*, 2013

## Excluded zones



To select zones of the image to discard in further calculations, you can:

- Define a threshold such as if  $Y < \text{Threshold}$  the pixel will be excluded
- Define a margin for the image (pixels on the borders to be excluded)
- Define a detector radius in case the detection area is circular.
- Add beamstop shaped excluded zone by giving a radius of the central part of the beamstop, a pixel inside the arm of the beamstop and the width of the arm.
- Add a polygonal excluded zone click ADD and click several points to define the zone.
- Add an arc-shaped excluded zone by clicking 3 points to define the zone in the following order: center, half radial width, half azimuthal aperture.
- Paint with the mouse the zones you want to exclude by clicking Mouse Free Paint and left-click and drag with the mouse. You can select the size of the square-shaped "brush" and undo the last change with the button if desired (or clear all).

After defining excluded zones, you may do one of the following:

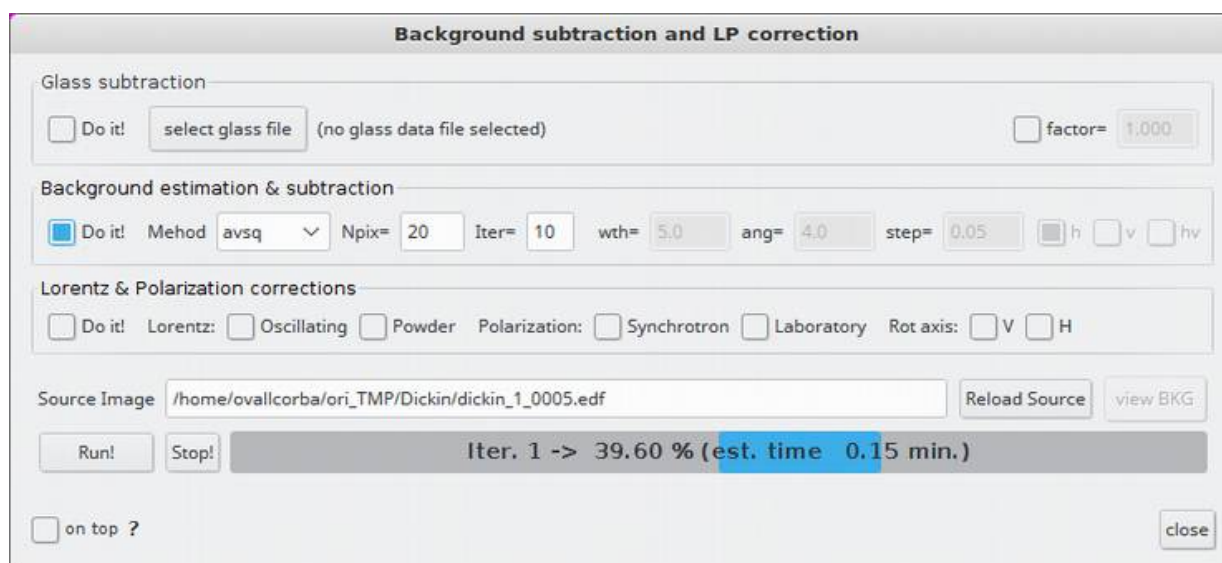
- Save an Excluded Zones (ExZ) file to be loaded later and/or to apply the zones to other images.
- Save a MASK image, which is an image in BIN format with all the intensities at zero

except for the mask pixels which have intensity -1.

- Save in a format (D2D, BIN) that contain the information.

On the main program window there is an option to show/hide the excluded zones, which are painted in magenta (paint ExZ) if activated.

## Background subtraction



There are 3 sections:

1. In the first one you can subtract a “glass” (or background) file by selecting it. A factor can be given (otherwise will be calculated in a conservative way and you can adjust it in next runs)
2. In the second one there are 5 methods to estimate the background, in summary:
  - **avsq**: Each iteration estimates the background by averaging square areas around each pixel from the previous iteration. Set the number of pixels for the side of the square (**Npix**) and the number of iterations (**Niter**). It is a slow process for high **Npix** and **Niter** values.
  - **avarc**: The same as previous option but using arc shaped areas (within 2-theta) around each pixel. Set the number of iterations (**Niter**) and the factors for the width (**wdt**) and angular aperture (**ang**) for the arcs. This is a very slow method.
  - **avcirc**: The background estimation for each pixel is the mean intensity from a radial integration (in the 2-theta circle containing each pixel). Set the stepsize



for the 2-theta ranges (step).

- **minsq**: The background intensity value for each pixel (v0) is calculated as: Minimum (v0, v1, v2, v3) where v1, v2 and v3 are related pixels applying a reflection of the image (vertical, horizontal and both). Set which operations to use (v,h,vh), and the number of pixels (Npix) defining the square zone to be averaged after the operation (use 0 to consider only 1 pixel). It is a fast method but some peak intensity may be subtracted.
  - **minarc**: The same as minsq but using an arc shaped zone for each pixel. Set the operations (v,h,vh) and the factors for width and angular aperture (wdt,ang).
3. The third one is to apply the corrections for Lorentz & polarization to the image pixels. Select the proper conditions (single grain oscillating/powder, synchrotron/lab, oscillating axis horizontal/vertical).

When clicking on Run! the sections marked with the **do it** tick will be executed.

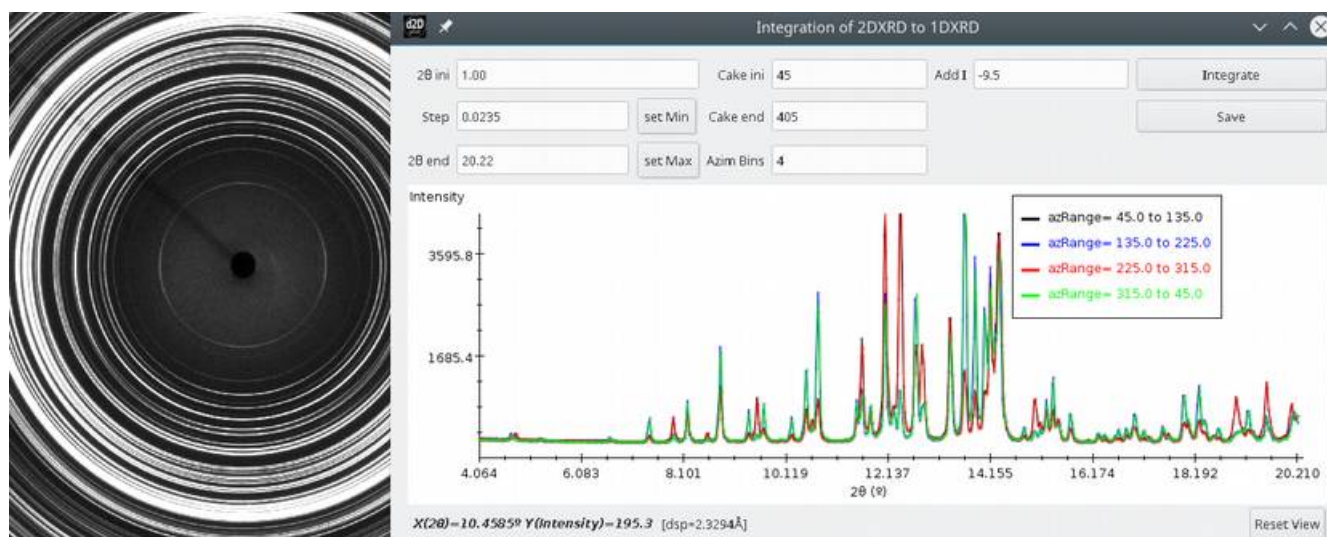
Visual inspection for residual peak intensity in the subtracted background can be done by clicking the [view BKG] button. Result images can be seen on the main window and source image can be reloaded if wanted. It is recommended to save the result to an image file before applying more corrections to the result file.

To subtract the background it is very important to define the excluded zones before.

## Integration of 2DXRD to 1DXRD

It performs the conversion of the 2D diffraction image to the 1D powder diffraction pattern given a 2-theta range and conditions (fields are self-explanatory). Result can be saved in a two columns file (2-theta intensity). Considerations:

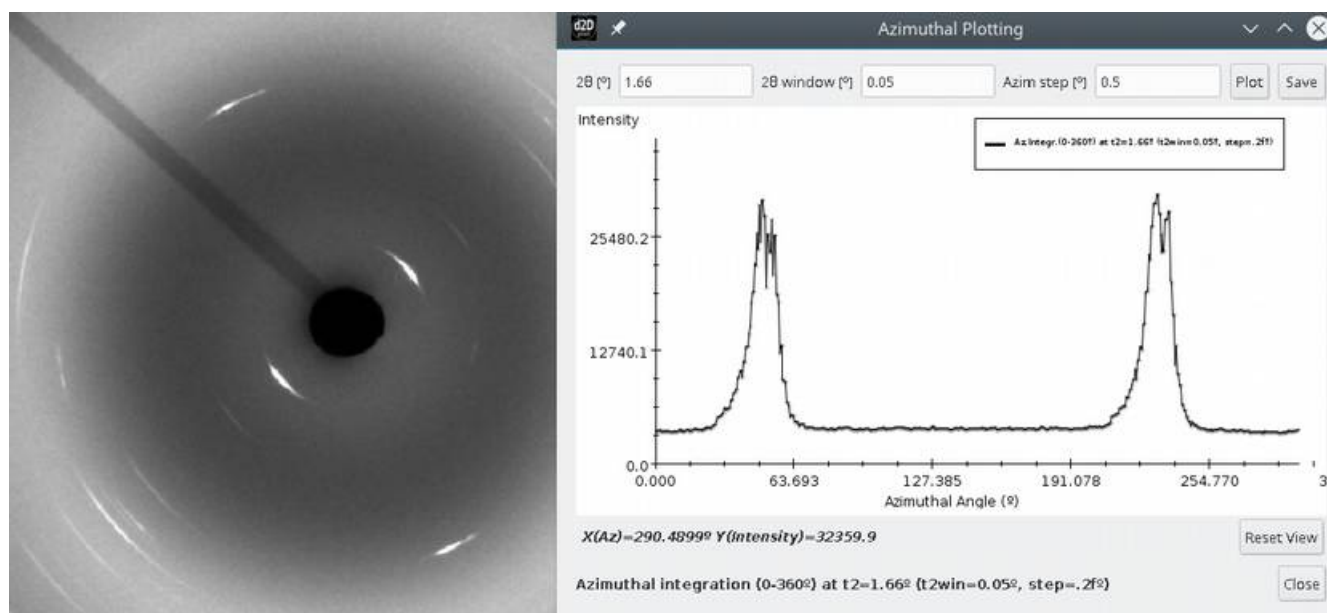
- The cake ini/end units are degrees and starting from the vertical (12h. on a clock) going positive clockwise.
- Azim bins (usually will be 1 for the full integration) is to divide the full integration (caki ini/end) in different cake fragments, so one pattern is generated for each part (check image above). It may be used to check for homogeneity or texture.
- Add I is to add intensity to the pixels of the image. It is useful in case the detector by default adds intensity to avoid zero. To subtract this extra intensity we need to introduce a negative value here.



Information for the radial integration methodology and geometrical corrections can be found on Hinrichsen, Dinnebier & Jansen, 2008; as the methodology implemented in d2Dplot follows the definitions on the book.

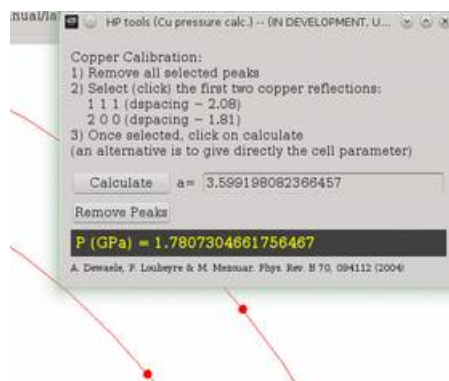
### Azimuthal (circular) plot

It performs a plot along the Debye ring specified by a 2-theta value with a tolerance (2-theta window) and an angular step (azim step). It is useful to check for graininess and for texture in powder samples.





## Copper pressure calculator (for High Pressure experiments)

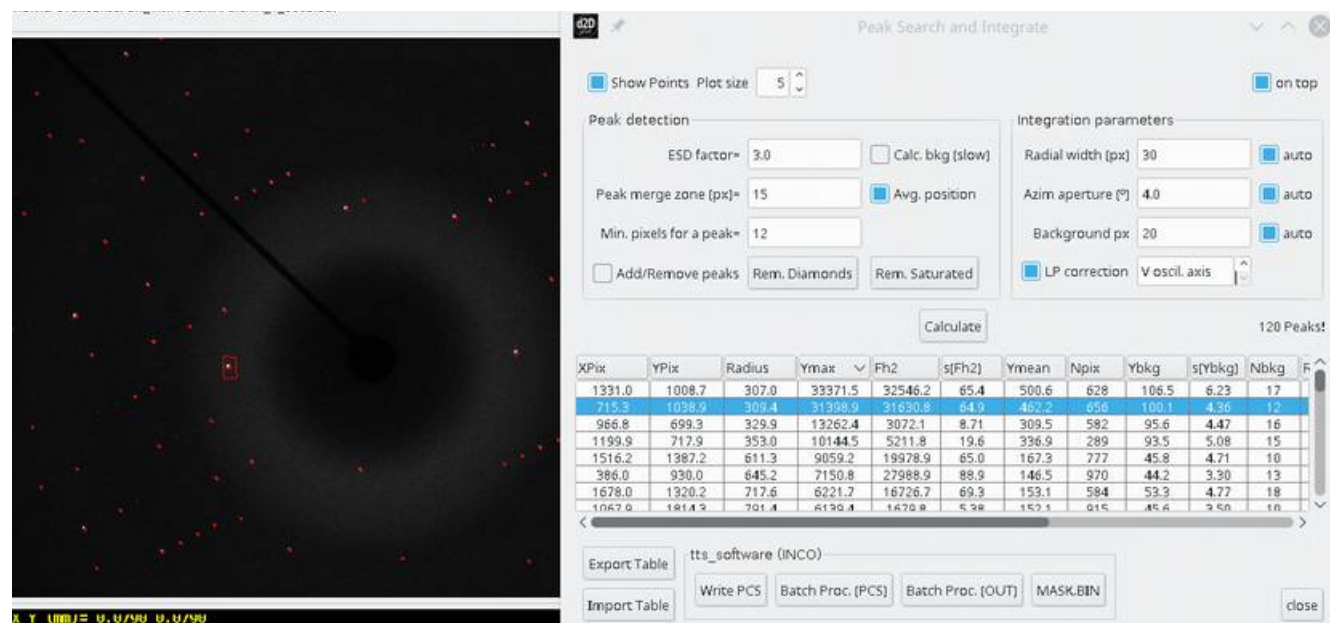


This module is still very preliminary and the only function it has implemented is the calculation of the pressure by the selection of two copper peaks (or alternatively directly from a given cell parameter).

## 4. Grain Analysis module

This module contains tools to work with images of single/few grains, but not powders.

### Find/Integrate Peaks



This is an important module of *d2Dplot* for the *TTS\_software* interaction. It finds and integrates spots on the image. The options for the peak detection are:

- ESD factor: It acts as a threshold related to the intensities standard deviation and optionally it can estimate the background for a better detection (it is slower).
- Peak merge zone: to avoid very close peaks. Avg. position means that when the merging is done it is taking the mass center of the peaks as the final peak position.
- Minimum number of pixels for a peak.
- By checking add/remove peaks, additional peaks can be added or removed by clicking with the left or right mouse button respectively. They are automatically integrated.
- Remove Diamonds tries to detect and remove the peaks coming from diamonds in case of Diamond Anvil Cells.
- Remove Saturated removes the saturated spots from the list.

The integration options are:

- Radial width of the integration zone (in pixels). Can be set to auto.

- Azimuthal aperture of the integration zone (in degrees). Can be set to auto.
- Number of pixels to calculate the background. Can be set to auto.
- Lorentz correction according to the oscillation axis.

The results are shown on a table with a lot of information:

- XPix, YPix = Pixel coordinates.
- Radius = Center to pixel vector modulus.
- Ymax = Maximum intensity.
- Fh2, s(FH2) = Integrated intensity and the associated standard deviation.
- Ymean = Mean intensity.
- Npix = Number of contributing pixels
- Ybkg, sYbkg = Background intensity that has been subtracted and its standard deviation.
- Nbkg = Number of background pixels used for the background estimation.
- RadWth = Radial width in pixels of the integration area.
- AzimDeg = Azimuthal aperture in degrees of the integration area.
- dsp = d-spacing of the spot.
- $p = \pi \cdot (y_{\max}/y_{\text{int}})^{2/3}$
- Swarm = If the peak has more than one maximum of intensity (may be overlap?)
- Satur = If the peak contain saturated pixels (indicates the number).
- nearMsk = True if the peak is close to a mask zone.

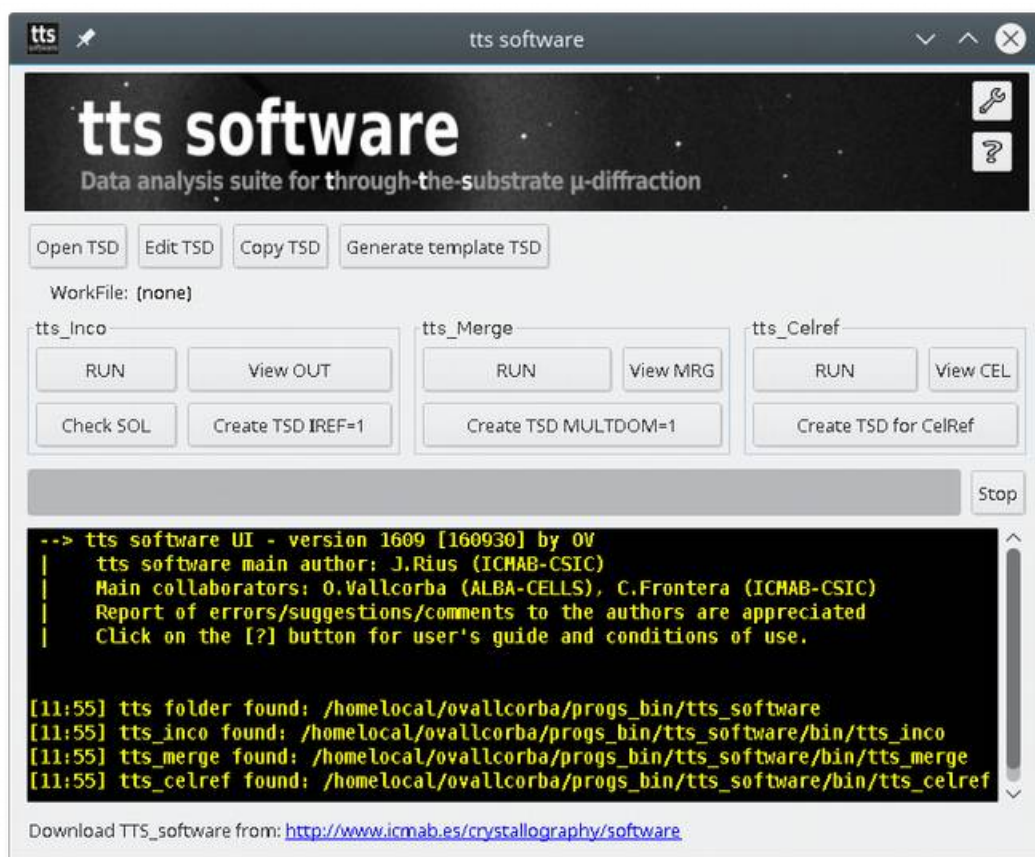
By clicking on a peak of the list the integration zone is shown on the image.

The peak list and intensity info can be exported as:

- A text file containing all the information of the table (it can be imported back).
- PCS file to be used in *tts-INCO* program (Rius *et. al.* 2015, 2016).

With the batch button several images can be processed using the same peak detection and integration parameters. The batch PCS generates an individual PCS file for each of the processed images. The batch OUT generates only one file containing the information of all the processed images. If the image on the main frame is changed (for example with the arrows on the top to navigate images), the peak search module is not closed and can be used to integrate directly the new image.

## Run TTS\_software



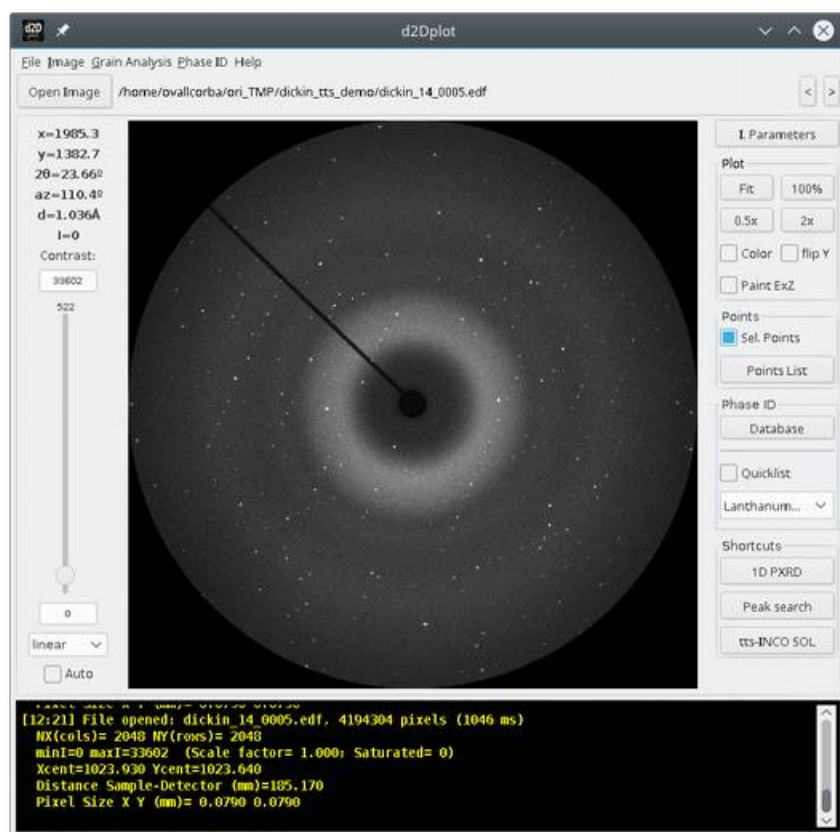
This option opens the front-end to the *TTS\_software*. This software is not included in *d2Dplot* and it should be downloaded separately from the website <http://departments.icmab.es/crystallography/software>. The first time we need to tell *d2Dplot* where the *TTS\_software* folder is located. It is done from the small configuration window that can be opened from the wrench tool button on the top right. The *TTS\_software* comes with its own manual but in this section you can find a small step-to-step tutorial for its application using the front-end.

### 1. Preparation of the Data

In order to apply the *TTS\_software* to extract and merge the intensities from one or several diffracting grains of a sample, first we need the following files in the working folder:

- Diffraction images with the proper nomenclature, which means a base name with a sequential numbering. In our case they will be: dickin\_14\_0000.edf, dickin\_14\_0001.edf, ... dickin\_14\_0010.edf. In total 11 images collected rotating  $\phi$  from -25 to +25° using  $\Delta\phi = 5^\circ$ .

- The PCS file for each one of the images, generated in the find/integrate peaks module explained in the previous section.
- A MASK.BIN file with the excluded zones (e.g. beam stop), generated in the Excluded Zones module explained in a previous section of this document.
- We open the central frame in *d2Dplot*, in our case dickin\_14\_0005.edf



## 2. Generate the input file

- In the *TTS\_software* frontend we click on **Generate template TSD** and we create the file `dickin_14.tsd` and start editing it after creation (the program will ask us).
- In the TSD file we must introduce fields **CELL**, **LATTICE** and **LAUE** according to our crystal. Then we set the **SWING** (which is  $\Delta\phi$ ), **DSFOU** (min *d*-spacing to be considered) and the option **IOFF=0** to indicate that we are performing a full scan of the central frame. **MULTDOM**, **ALON**, **ALAT**, **ASPIN** are left at zero at this stage as they are not used. For **MODEL** we also put zero (as none is used) and for the **PCS** we put 1 file which is the number 5 with the offset 0°. **NSOL** is the number of solutions that the program will output. At the end it should look like this:

```

DICKINSONITE C2/c
CELL
  16.625 10.0746 24.8365 90.0 105.24 90.0
LATTICE
C
LAUE
2
&CONTROL
SWING=5.0,
DSFOU=1.0,
MULTDOM=0,
IOFF=0,
NSOL=10,
ALON=0.0,
ALAT=0.0,
SPIN=0.0
/
MODEL
  0
PCS/HKL
  1
5, 0.0

```

- We save and close the TSD file and click RUN on the tts\_inco section. The output will show something like this:

```

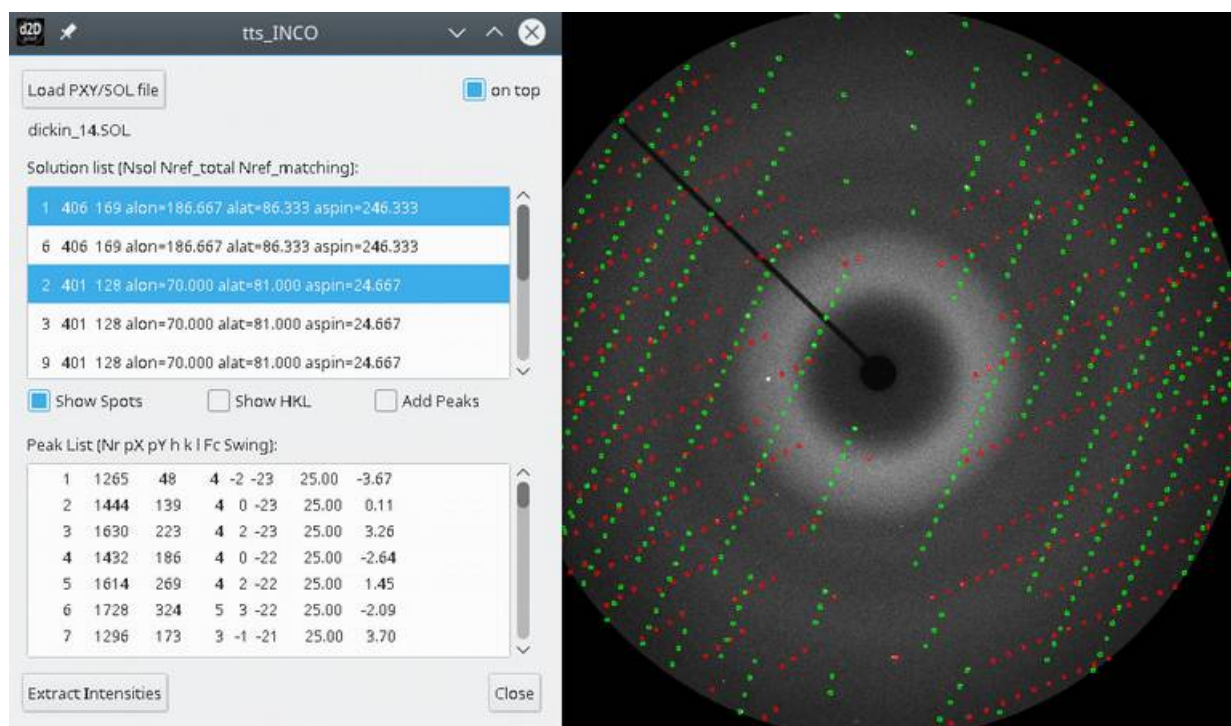
[12:54] -----
[12:54] NPIX X, Y, SKF, CENT X, CENT Y=      2048      2048      1.00000000      1023.92999      1023.64001
[12:54] PIXLEN, SEP OD, WAVE      79.0000000      185.169998      0.42460005
[12:54] DSCIRC, DSFOU:      1.03819311      1.03819311
[12:54] 0/1/3 ES ELECTRONS/RAIGS_X/WOTT:      1
[12:54] DELTA PHI -> EIX GIR VERTICAL:      5.00000000
[12:54] MU & GRUIX VIDRE:      3.39000010      0.00000000
[12:54] NUM. TEORIC DE REFLEXIONS:      417
[12:54]
[12:54] NOW DEL PCS: dickin_14.0005
[12:54] ANGLE DE OFFSET en graus:      0.00000000
[12:54] NPICS (.PCS) =      307
[12:54] IHMAX, IKMAX, ILMAX:      16      10      24
[12:54] NCIRC:      7548
[12:54] SCAN COMPLET:
[12:54] NGRA=      1      188.000000      84.000000      247.000000      414      108      55.8730659
[12:54] NGRA=      2      70.000000      82.000000      24.000000      412      107      53.5932808
[12:54] NGRA=      3      71.000000      78.000000      25.000000      407      98      49.3191910
[12:54] NGRA=      4      6.00000000      88.000000      246.000000      408      79      56.9440575
[12:54] NGRA=      5      174.000000      92.000000      66.000000      408      79      56.9440498
[12:54] NGRA=      6      182.000000      86.000000      246.000000      415      79      56.0348167
[12:54] NGRA=      7      8.00000000      77.000000      248.000000      422      74      55.2485123
[12:54] NGRA=      8      8.00000000      84.000000      247.000000      413      70      56.4288565
[12:54] NGRA=      9      73.000000      81.000000      25.000000      431      70      51.9858589
[12:54] NGRA=     10      2.00000000      79.000000      247.000000      414      67      58.7453766
[12:54]
[12:54] NGRA=      1      186.666672      86.3333282      246.333252      406      109      52.4905396
[12:54] NGRA=      2      70.000000      81.000000      24.6666088      402      128      51.7998543
[12:54] NGRA=      3      70.000000      81.000000      24.6666183      402      128      51.7998543
[12:54] NGRA=      4      6.66666603      86.3333359      246.333237      408      93      54.4115219
[12:54] NGRA=      5      173.333328      93.6666641      66.3332443      408      93      54.4115067
[12:55] NGRA=      6      186.666656      86.3333359      246.333237      406      109      52.4905396
[12:55] NGRA=      7      5.00000286      78.6666641      247.333252      425      81      55.1918526
[12:55] NGRA=      8      6.66666794      86.3333282      246.333252      408      93      54.4115219
[12:55] NGRA=      9      70.000000      81.000000      24.6666183      402      128      51.7998543
[12:55] NGRA=     10      6.99999475      84.9999924      246.333252      405      85      55.2678947
[12:55]
[12:55]
[12:55] INCO finished correctly.

```



### 3. Checking the results

Now we can click on check SOL to see on the diffraction image the matching of the orientations found with the diffraction peaks. We can select the different solutions on the list to see the reflections on the image. In this case we can see that there are two different solutions for orientations that match the spots on the image (with 169 and 128 spots perfectly matching respectively).



We can also check the OUT file where at the end there is a matrix with the number of coincident peaks between solutions, where we can clearly see that the two solutions correspond to different grains, as they do not share the same peaks.

### 4. Extraction of partial hkl datasets from one oriented domain

Now we have two domains oriented. We can select the first one (solution 1) and click on the Create TSD IREF=1 button. When the program asks for a file name we put the same as before but adding "d1" to identify that it corresponds to the domain number 1 (so it will be dickin\_14d1.tsd). We can check the contents of the file but it is not necessary to edit anything as the program already populated the list of images with the corresponding offsets, the IOFF=1 (to indicate that now we are exploring a complete dataset with one orientation set) and the ALON, ALAT and SPIN angles according to the orientation selected as solution.

```

DICKINSONITE C2/c
CELL
  16.625 10.0746 24.8365 90.0 105.24 90.0
LATTICE
C
LAUE
2
&CONTROL
SWING=5.,
DSFOU=1.0,
MULTDOM=0,
IOFF=1,
NSOL=10,
ALON=186.667,
ALAT=86.333,
SPIN=246.333,
/
MODEL
  0
PCS/HKL
11
0, -25.00
1, -20.00
2, -15.00
3, -10.00
4, -5.00
5, 0.00
6, 5.00
7, 10.00
8, 15.00
9, 20.00

```

The WorkFile now should be the new TSD file. We can RUN again inco. Now the output is slightly different as it went through all the images extracting the peaks according to the orientation entered and the offset angle of the image.



```

[13:16] IHMAX,IKHAX,ILMAX:      16      10      24
[13:16] NCIRC:      7548
[13:16] REFINEMENT OF SOLUTION:
[13:17] NGRA=      1  186.667007      86.3330002      246.332901      437      191  48.6158104
[13:17]
[13:17] NOM DEL PCS: dickin_14_0007
[13:17] ANGLE DE OFFSET en graus:  10.0000000
[13:17] NPICS (.PCS) =      309
[13:17] IHMAX,IKHAX,ILMAX:      16      10      24
[13:17] NCIRC:      7548
[13:17] REFINEMENT OF SOLUTION:
[13:17] NGRA=      1  186.667007      85.9996643      246.666229      424      167  46.3864670
[13:17]
[13:17] NOM DEL PCS: dickin_14_0008
[13:17] ANGLE DE OFFSET en graus:  15.0000000
[13:17] NPICS (.PCS) =      303
[13:17] IHMAX,IKHAX,ILMAX:      16      10      24
[13:17] NCIRC:      7548
[13:17] REFINEMENT OF SOLUTION:
[13:17] NGRA=      1  186.333679      86.3330002      246.666229      396      143  49.9993019
[13:17]
[13:17] NOM DEL PCS: dickin_14_0009
[13:17] ANGLE DE OFFSET en graus:  20.0000000
[13:17] NPICS (.PCS) =      319
[13:17] IHMAX,IKHAX,ILMAX:      16      10      24
[13:17] NCIRC:      7548
[13:17] REFINEMENT OF SOLUTION:
[13:17] NGRA=      1  186.333679      85.9996643      246.666229      406      155  50.6143227
[13:17]
[13:17] NOM DEL PCS: dickin_14_0010
[13:17] ANGLE DE OFFSET en graus:  25.0000000
[13:17] NPICS (.PCS) =      296
[13:17] IHMAX,IKHAX,ILMAX:      16      10      24
[13:17] NCIRC:      7548
[13:17] REFINEMENT OF SOLUTION:
[13:17] NGRA=      1  186.333679      85.6663361      246.999557      397      155  51.4432335
[13:17]
[13:17] INCO finished correctly.

```

At this point, we should have at the folder the partial HKL files (and SOL files) for each of the images, considering only the domain 1.

### 5. Merge the partial dataset of one domain

To merge the datasets of one domain the input file is the same TSD used in the previous step, so we click directly the RUN button in the tts\_Merge section. The output will be similar to this one:

```

[13:22] =====
[13:22] TTS_MERGE - MERGING OF PARTIAL/MULTIDOMAIN HKL FILES -
[13:22] Copyright (C) Jordi Rius (ICMAB-CSIC) v.160928
[13:22] Institut de Ciencia de Materials de Barcelona (CSIC)
[13:22] 08193-Bellaterra, Catalonia, Spain
[13:22] Lahey/Fujitsu Fortran 95 Compiler
[13:22] =====
[13:22] NUM. IMATGES:      11
[13:22] ILAUE:      2 DSFOU:      1.00000000
[13:22] NIMAG      11
[13:22] DSMIN:      1.03982651
[13:22] RESIDUAL INICIAL:      1  4.43693995E-02
[13:22] NRFASI      2262
[13:22] NSYS=      1
[13:22] RESIDUAL FINAL:      4.36209030E-02
[13:22] FACTORS ESCALA (F"S):
[13:22] 1.01212800      1.03857911      0.997871697      1.00442052      1.00212157      0.994581044      0.98775
[13:22] N. REFL. IN ASYMMETRICAL UNIT (DOWN TO DSFOU):      2262
[13:22] N. MISSING REFLECTIONS:      1782
[13:22] OBSERVED REFLECTIONS(%):      21.2201595
[13:22] MERGE finished correctly.

```

where we can see the scale factors applied to each of the partial datasets and the residual of the merge process. Checking the residual and the evolution of the scale

factor is a good indication of the consistency of the data. This information and a summary of the process can also be checked in the MRG file.

The result is an HKL file (dickin\_14d1.HKL), containing the extracted intensities for this domain.

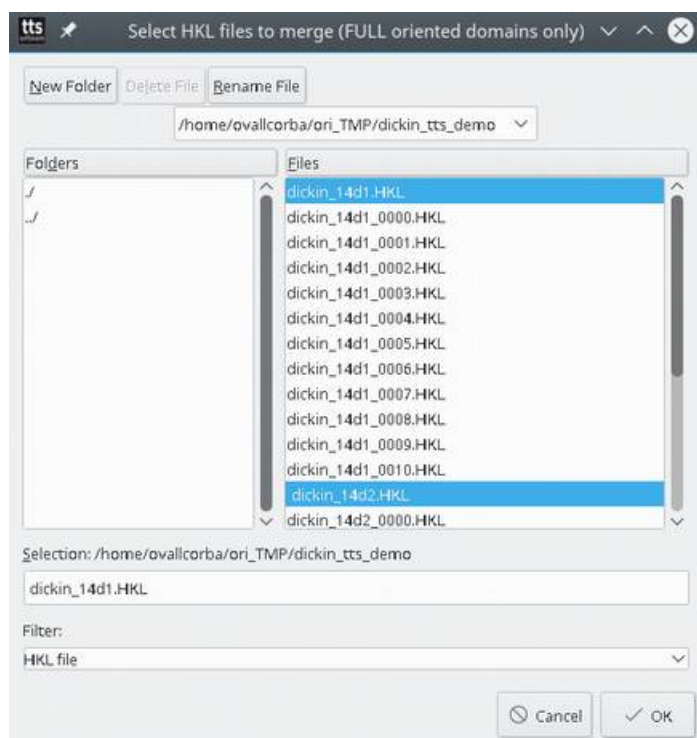
## 6. Extract the intensities for all the different domains

Steps 4 and 5 should be repeated for the different domains contained on the same set of images. Also, it may be the case that we have other sets of images containing different oriented grains of the same compound so we should repeat the process for them in order to get all the HKLs for the different individual domains.

In this example we repeat steps 4 and 5 using the solution number 2, which is the orientation found for the second domain on our set of images.

## 7. Multidomain merging

Finally we will perform the merging of datasets from multiple oriented domains. In this example we have datasets (HKL files) from two domains: dickin\_14d1.HKL and dickin\_14d2.HKL. To perform the multidomain merging we have to click on Create TSD MULTDOM=1. Here the program will ask to save a new TSD file, which in this case the name has no restrictions. Then the program will ask to select the HKL files we want to merge and we have to select the two corresponding to the full oriented datasets:



8. We do not need to change anything from the new TSD file created, it will have MULTDOM=1 and the HKL files:

```
DICKINSONITE C2/c
CELL
  16.625 10.0746 24.8365 90.0 105.24 90.0
LATTICE
C
LAUE
2
&CONTROL
SWING=5.,
DSFOU=1.0,
MULTDOM=1,
```

(...)

```
PCS/HKL
2
dickin_14d1.HKL
dickin_14d2.HKL
```

We RUN `tts_Merge` again with this new TSD file and the output will be similar as the merging of partial datasets (with the merge residual and scale factors).

```
[14:10] =====
[14:10] TTS_MERGE - MERGING OF PARTIAL/MULTIDOMAIN HKL_FILES -
[14:10] Copyright (C) Jordi Rius (ICMAB-CSIC) v.160928
[14:10] Institut de Ciencia de Materials de Barcelona (CSIC)
[14:10] 08193-Bellaterra, Catalonia, Spain
[14:10] Lahey/Fujitsu Fortran 95 Compiler
[14:10] =====
[14:10] NUM. IMATGES:          2
[14:10] ILAUE:          2 DSFOU:    1.00000000
[14:10] NIMAG          2
[14:10] DSMIN:    1.03982651
[14:10] RESIDUAL INICIAL:          1  4.75926995E-02
[14:10] NRFAI          2262
[14:10] NSYS=          1
[14:10] RESIDUAL FINAL:    4.75926995E-02
[14:10] FACTORS ESCALA (F"S):
[14:10]    1.01421189    0.985583127
[14:10] N. REFL. IN ASYMMETRICAL UNIT (DOWN TO DSFOU):    2262
[14:10] N. MISSING REFLEXIONS:    1454
[14:10] OBSERVED REFLECTIONS(%):    35.7206001
[14:10] MERGE finished correctly.
```

At the working folder, the final HKL file (with the same name as this last TSD file) containing the intensity information of all the merged data will be generated in the SHELX HKLF 4format (Sheldrick, 2014), to be used for structure refinement.

As an example of the complexity of the refinements that can be performed, the SHELX output file .RES in space group C2/c of the dickinsonite phosphate from Cuevas (Argentina), derived from 3 domains with 11 images each one, is given below. The ttr-diffraction data were measured at ALBA Synchrotron (Barcelona, Spain) in collaboration with Dr. Fernando Colombo (CICTERRA-CONICET, Córdoba, Argentina) in the frame of Project MAT2015-67593-P of MINECO & FEDER. The formula for this dickinsonite sample is  $\text{K}_{0.5}\text{Li}_{0.2}\text{Na}_{5.4}\text{Ca}_{0.9}(\text{Mn}_{9.8}\text{Fe}_{4.0})\text{Al}_{0.8}(\text{PO}_4)_{12}(\text{OH})_2$

```

TITL  DICKINSONITE C2/c K0.5Li.2Na5.4Ca0.9(Mn9.8Fe4.0)Al0.8(P04)12(OH)2
CELL  0.43460  16.62506  10.0747 24.8365  90.000  105.24066  90.000
ZERR  1          0.001    0.001    0.001    0.0    0.01    0.0
LATT  7
SYMM  X, -Y, Z+1/2
SFAC  MN  CA  FE  P  AL  NA  O  K  F
UNIT  56   4   4  48   4  12 200  1  1
L.S.  100
FMAP  2
BOND  0.5
PLAN  25
LIST  5
DFIX  51  0.01  P1 01A  P1 01B  P1 01C  P1 01D
DFIX  51  0.01  P2 02A  P2 02B  P2 02C  P2 02D
DFIX  51  0.01  P3 03A  P3 03B  P3 03C  P3 03D
DFIX  51  0.01  P4 04A  P4 04B  P4 04C  P4 04D
DFIX  51  0.01  P5 05A  P5 05B  P5 05C  P5 05D
DFIX  51  0.01  P6 06A  P6 06B  P6 06C  P6 06D
DFIX  51.633  0.05 01A 01B 01A 01C 01A 01D 01B 01C 01B 01D 01C 01D
DFIX  51.633  0.05 02A 02B 02A 02C 02A 02D 02B 02C 02B 02D 02C 02D
DFIX  51.633  0.05 03A 03B 03A 03C 03A 03D 03B 03C 03B 03D 03C 03D
DFIX  51.633  0.05 04A 04B 04A 04C 04A 04D 04B 04C 04B 04D 04C 04D
DFIX  51.633  0.05 05A 05B 05A 05C 05A 05D 05B 05C 05B 05D 05C 05D
DFIX  51.633  0.05 06A 06B 06A 06C 06A 06D 06B 06C 06B 06D 06C 06D
WGHT  0.133500 3082.684570
FVAR  0.05862  0.01182  0.01200  0.00788  1.55710
NA1   6   10.000000  10.500000  10.000000  10.50000  0.02762
NA2   6   10.250000  10.250000  10.000000  10.50000  0.02420
K3    8   10.000000  0.005378  10.250000  10.50000  0.00001
NA4   6   0.132936  1.015633  0.119276  0.67743  41.00000
OH1   9   0.233381  0.503695  0.139281  11.00000  21.00000
AL1   5   10.000000  10.000000  10.000000  10.50000  21.00000
P1    4   0.101234  0.266482  0.048055  11.00000  21.00000

```

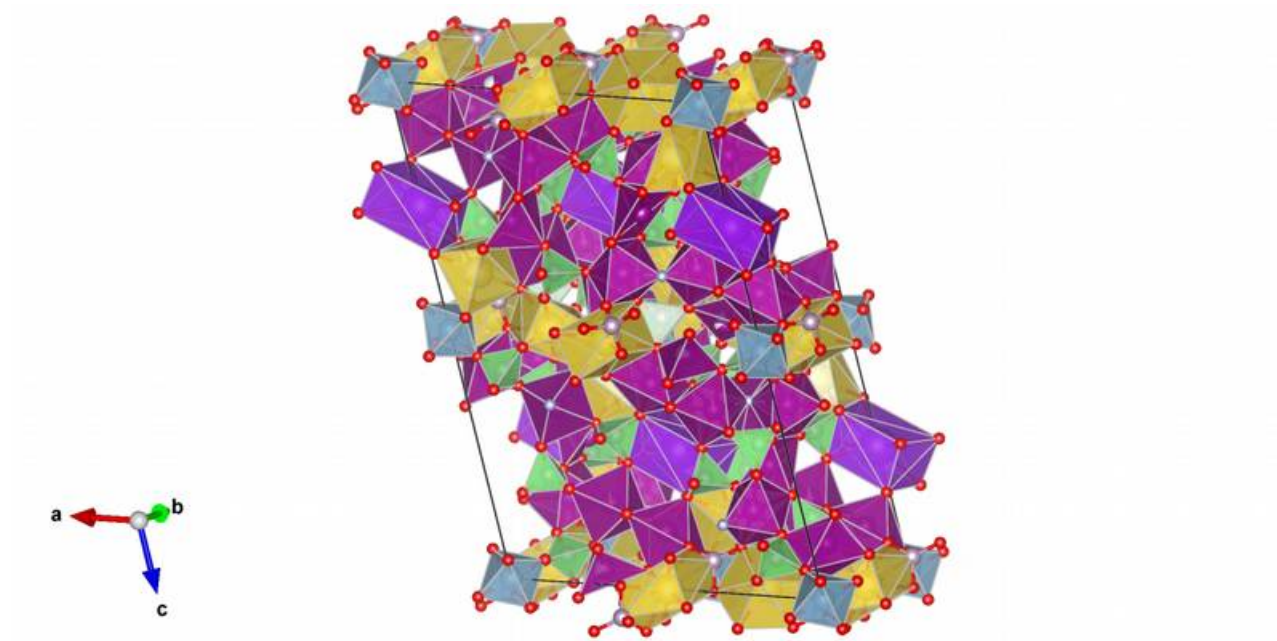
01A	7	0.046363	0.140677	0.046868	11.00000	21.00000
01B	7	0.065798	0.375386	0.080061	11.00000	21.00000
01C	7	0.094964	0.318291	-0.011999	11.00000	21.00000
01D	7	0.194034	0.236335	0.077738	11.00000	21.00000
P2	4	0.124733	0.749571	0.037486	11.00000	21.00000
02A	7	0.104062	0.687555	-0.021989	11.00000	21.00000
02B	7	0.221242	0.751802	0.060897	11.00000	21.00000
02C	7	0.083064	0.665482	0.075269	11.00000	21.00000
02D	7	0.092371	0.894892	0.035196	11.00000	21.00000
P3	4	0.369061	0.466728	0.076381	11.00000	21.00000
03A	7	0.447457	0.449182	0.053776	11.00000	21.00000
03B	7	0.368578	0.345043	0.114996	11.00000	21.00000
03C	7	0.376093	0.599224	0.109382	11.00000	21.00000
03D	7	0.287863	0.465437	0.027768	11.00000	21.00000
P4	4	0.127794	0.225436	0.211515	11.00000	21.00000
04A	7	0.126217	0.174766	0.270341	11.00000	21.00000
04B	7	0.201643	0.158136	0.194396	11.00000	21.00000
04C	7	0.134268	0.379504	0.210496	11.00000	21.00000
04D	7	0.044440	0.183265	0.169512	11.00000	21.00000
P5	4	0.355820	0.301319	0.296438	11.00000	21.00000
05A	7	0.339577	0.359069	0.236380	11.00000	21.00000
05B	7	0.441640	0.345068	0.334858	11.00000	21.00000
05C	7	0.283200	0.353624	0.319324	11.00000	21.00000
05D	7	0.351220	0.146694	0.294494	11.00000	21.00000
P6	4	0.383165	1.014769	0.131351	11.00000	21.00000
06A	7	0.472475	1.064962	0.132976	11.00000	21.00000
06B	7	0.385972	0.860560	0.137727	11.00000	21.00000
06C	7	0.353089	1.080269	0.179339	11.00000	21.00000
06D	7	0.318630	1.048330	0.075051	11.00000	21.00000
A1	1	0.019823	0.750835	0.402829	0.99006	41.00000
A2	1	0.105462	0.515578	0.139693	0.98645	41.00000
A3	1	0.465396	0.744221	0.100429	1.00000	41.00000
A5	1	0.214548	0.509123	0.267225	0.99424	41.00000
A6	1	0.295040	0.700975	0.147117	0.94509	41.00000
A7	1	0.280683	0.706379	0.658489	0.96677	41.00000
A8	1	0.219351	0.602265	-0.024549	0.63835	41.00000
HKLF	4					
REM		DICKINSONITE C2/c K0.5Li.2Na5.4Ca0.9(Mn9.8Fe4.0)Al0.8(P04)12(OH)2				
REM	R1 =	0.1331	for	907 Fo > 4sig(Fo)	and	0.1340 for all 916 data
REM		134	parameters refined using	60	restraints	
END						
WGHT		0.0861		3480.0105		



REM Highest difference peak 4.708, deepest hole -2.539, 1-sigma level 0.454

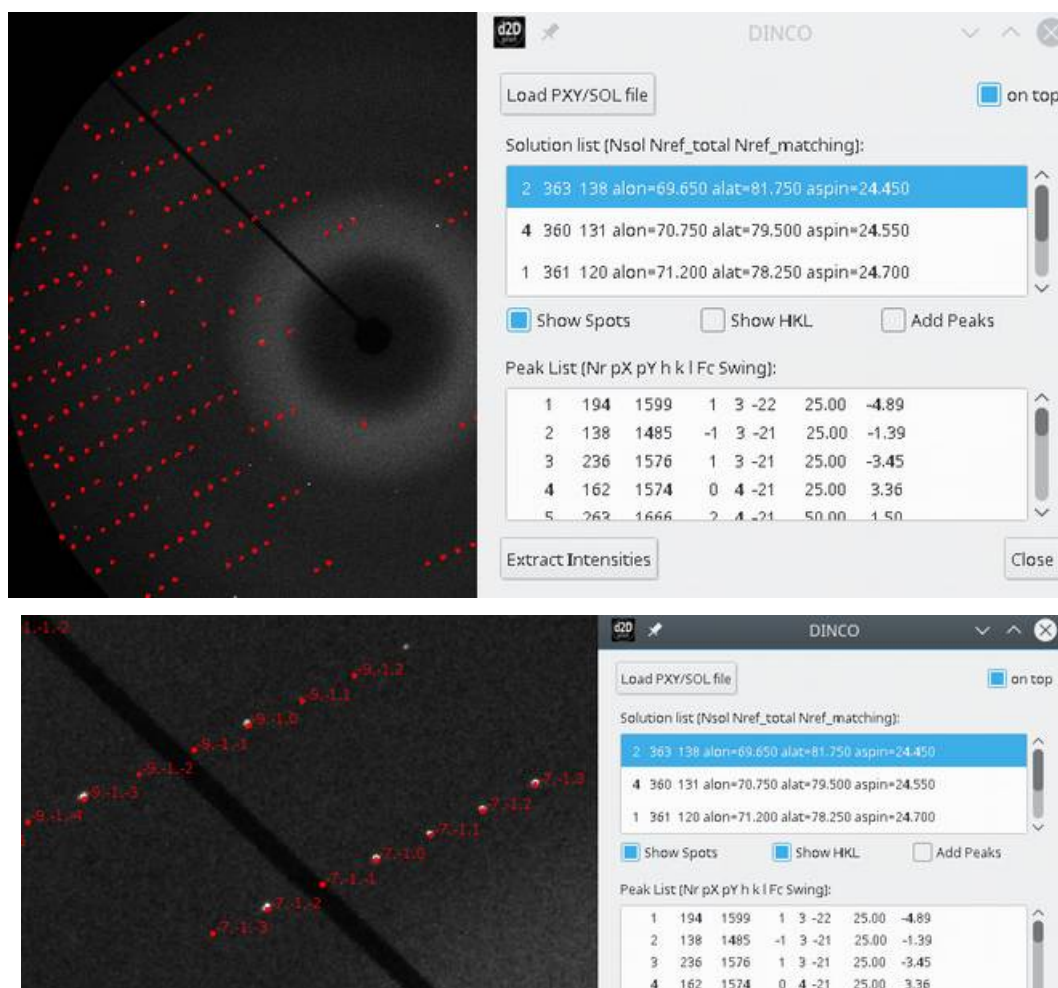
Q1	1	0.2244	0.2857	-0.0214	11.00000	0.05	4.71
Q2	1	0.5000	1.1384	0.2500	10.50000	0.05	2.97
Q3	1	0.4261	0.9968	0.1663	11.00000	0.05	2.69
Q4	1	0.0019	-0.0013	0.2731	11.00000	0.05	2.69
Q5	1	0.1778	1.0025	0.1109	11.00000	0.05	2.62
Q6	1	0.0000	0.4305	0.2500	10.50000	0.05	1.76
Q7	1	0.2347	0.7048	-0.0272	11.00000	0.05	1.58
Q8	1	0.0765	0.4784	0.2261	11.00000	0.05	1.55
Q9	1	0.2033	0.5138	-0.0425	11.00000	0.05	1.42

Perspective view of the refined unit cell contents of dickinsonite (refined in  $C2/c$ ). Polyhedra colors: P (green), Na (yellow), K (cyan), Al (blue) and Mn/Fe (violet). Removal the center of symmetry (space group  $Cc$ ) caused the  $R_1$  value to drop to 0.089.



## Load *tts-INCO* SOL/PCS files

Here the output files from *tts-INCO* can be opened to display the reflections on the image for the multiple solutions. If there are more than one, can be selected simultaneously (painted in different colors) to check if there are multiple grains. HKL indexes can be shown, peaks can be added (activating the option add peaks to click on the image) or removed (delete from list or by right-clicking).



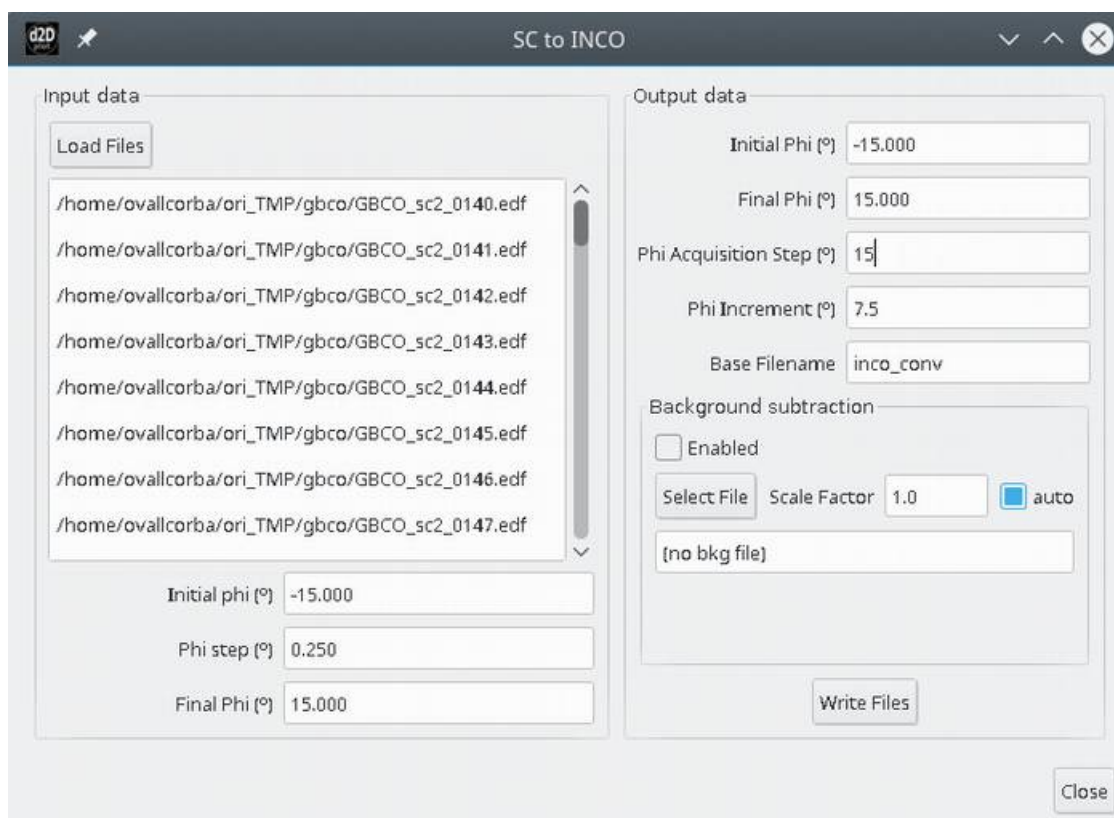
Also the peaks can be integrated by clicking **extract intensities** and a window equivalent as the one in **find/integrate peaks** will be opened.

## Load XDS file

This option is similar as the previous one but with a *spot.xds* file from *XDS* (**X**-ray **D**etector **S**oftware, CCP4; Kabsch, 1988).

## SC data to INCO

This option is used to sum image files specifying ranges of measurement.



For example, taking single crystal data collected from  $-15^\circ$  to  $15^\circ$  rotation with steps of  $0.25^\circ$  (120 images), it can be converted to 3 images of  $15^\circ$  rotation each with overlapping of  $7.5^\circ$ . In this case, 3 images will be generated: 1) data from  $-15$  to  $0^\circ$ ; 2) data from  $-7.5$  to  $7.5^\circ$ ; 3) data from  $0$  to  $15^\circ$ . Usually if you have single crystal data there is no need to perform any sum of the data, this is only intended in case `tts_INCO` wants to be used for any specific reason or to check data by performing other combinations of sums. Optionally, a background file can be subtracted to all the individual files before adding them up.



## 5. Phase ID

This is a strong part of *d2Dplot* which allows the fast identification of compounds from a custom database. There are two different “compound databases” considered in the program (actually lists will be the correct term to refer to them and not database). The full database, which is accessed via **Phase\_ID - Database** menu or the **Database** button in the main window, and the QuickList database which is intended to be a much smaller one (a selection of compounds from the full database) and that can be accessed directly from the main window. The “databases” (or lists) are stored in plain-text files and the paths for the default ones that are automatically considered by the program are given in the `d2dconfig.cfg` file.

### Database

**Compound DB**

Load DB Save DB ☒ ShowRings ? on top ☐

☐ Apply name filter:

Import CIF Import HKL

Name [Formula] (alt. names)

Silicon [Si]  
 Lanthanum hexaboride [La B6]  
 Diamond [C]  
 Graphite [C]  
 Nickel [Ni]  
 Nickel oxide [Ni O]  
 Tungsten [W]  
 Sodium chloride [Na Cl] (aka: Ha...  
 Cadmium chloride [Cd Cl2]  
 Aluminium oxide [Al2 O3] (aka: C...  
 Aluminium oxide (kappa) [Al2 O3]  
 Aluminium oxide (gamma) [Al2.666...]

Name Silicon  
 Name (alt)   
 Formula Si  
 Cell parameters '5.4307 90.000 90.000 90.000'  
 Space group Fd 3 m Calc Refl.  
 Reference ystal Structures 1 (1963) 7-83  
 Comment

list of (one per line): h k l d-spacing intensity

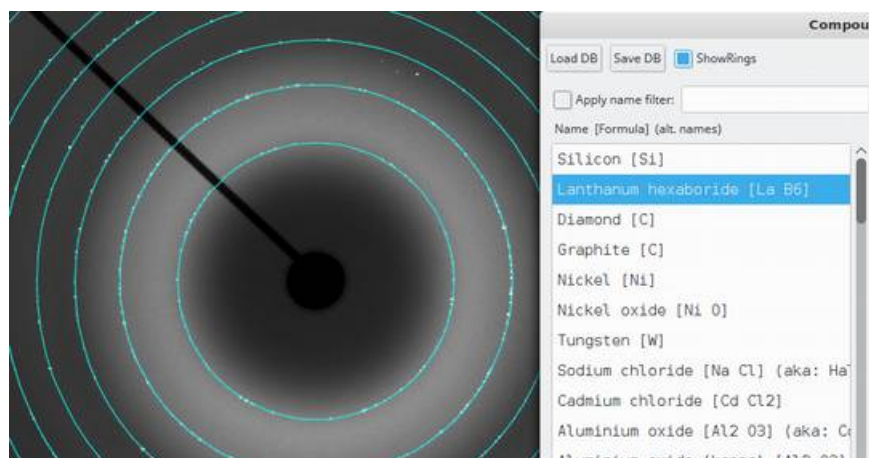
1	1	1	3.13540	100.00
2	2	0	1.92000	64.31
3	1	1	1.63740	37.32
4	0	0	1.35770	9.58
3	3	1	1.24590	14.14
4	2	2	1.10850	19.38

New Remove Search by peaks reset list Add to Quicklist Apply Changes Add as New

close

Here a plain-text DB file can be loaded. By default it opens the file `default.db` (which is in the program folder) as the example one coming with *d2Dplot*. Once loaded:

- Click on any compound to see the expected diffraction rings position on the image (if ShowRings is selected)



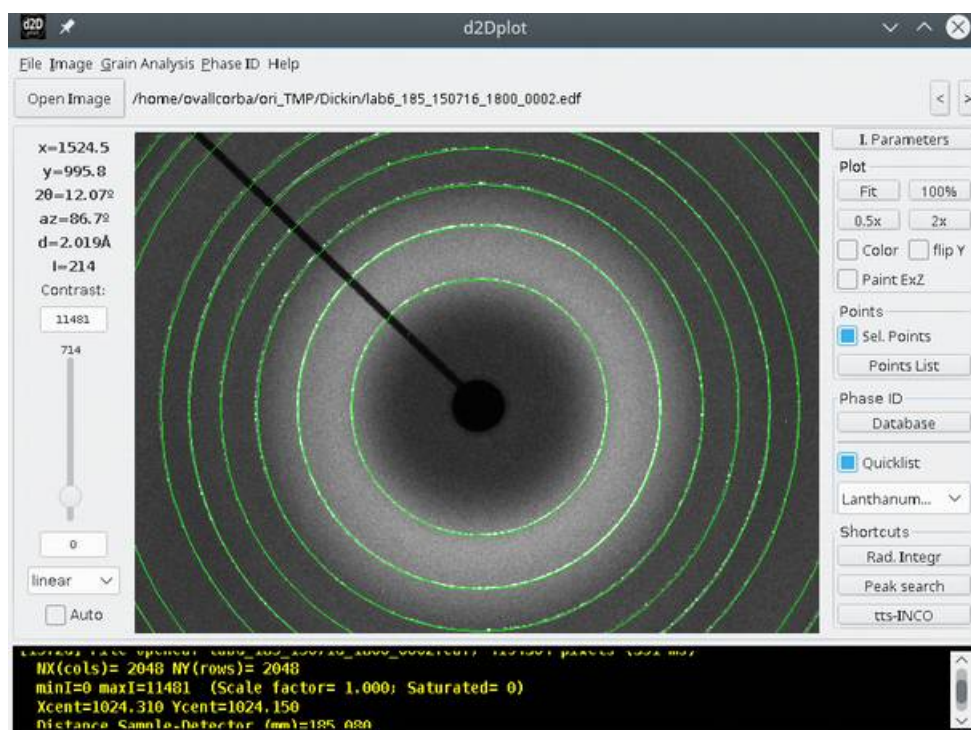
- Check apply name filter and type to easily find the desired compound
- Any selected compound may be edited from the fields on the right section of the window and clicking apply changes to update it or add as new to copy it as a new entry. Also new compounds may be added or removed by clicking new or remove. For new compound the information should be introduced. If the unit cell and space groups are known, the expected reflection positions can be calculated with calc Refl and the hkl list will be updated automatically. Alternatively, an HKL file or a CIF file can be imported. For CIF files, the hkl list (with calculated structure factors) will be automatically generated taking the cell parameters, symmetry and atom positions from the file. A confirmation window will show the information retrieved from the CIF to check for correctness.

Label	Atom Type	x/a	y/b	z/c	Occ	U_iso
Na1	Na1+	0	0	0	1	0
Cl1	Cl1-	0.5	0.5	0.5	1	0

OK Cancel

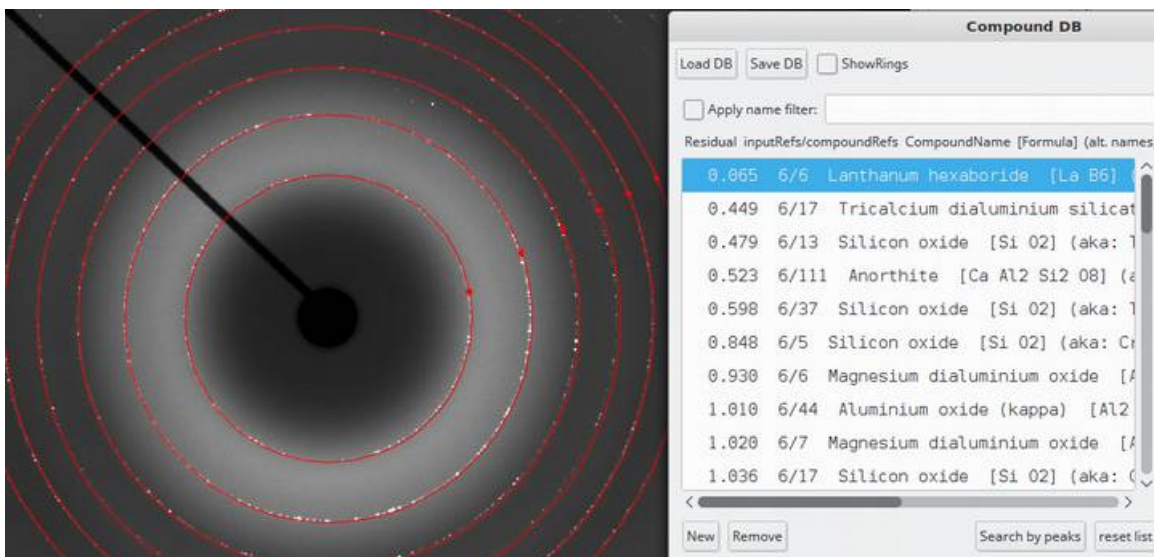
- Alternatively you can edit manually the DB file. It is a simple self-explanatory text file and its format is explained in another section of this guide.

- Add to QuickList (QL) to be able to access the rings from the main window directly. Compounds in the QL are saved in a separate file with the same format as the DB file and can also be edited the same way



There is the possibility to search in the database by image peak positions (selected on the main image window by the select points tool). To search by peaks:

- On the main window click on the desired rings so that they are selected in the point list (Sel.points should be active)
- Click the button **search by peaks**. There are two options that affect the result of the search:
  - To consider the total number of reflections up to the d-spacing corresponding to the last input ring (recommended).
  - To consider the intensity of the rings (not recommended, only if the rings are well defined, the database contain intensity info and the first option did not gave good results).
- List will be updated by the best matching compounds (with respective residuals)
- Click on the compounds to see the rings on top of your image and check if there is a good match.



**Note:** The purpose of this database system in *d2Dplot* is to allow you (the user of the program) creating your own database with your choice of compounds (e.g. the family of compounds you are working with as possible candidates for phase identification). There are several compound databases where you can find X-ray diffraction information, including *d*-spacings to introduce to your *d2Dplot* database. These databases can be proprietary ([ICDD](#), [ICSD](#), [CCDC](#),...) so that you need to purchase a license, or free ([COD](#)). The author of *d2Dplot* takes no responsibilities regarding where the final users of the program gets the X-ray diffraction information or its correctness.

The default DB is a small selection of 60 compounds taken from different sources. Each entry contains the reference from where it has been taken (with the respective authors) which can be retrieved from the reference field on each entry of the database. If any of these entries should be removed (for whatever reason) please contact the author and they will be removed immediately.

## 6. Macro mode

This is a very recent feature of the program. It is the way to make the program do some operations to images and get directly the output, without opening any GUI if possible. The *macro* mode can be accessed via the command line. It is accessed by entering *-macro* as first argument when launching the program from the command line followed by the filename to the image to be processed. Then following arguments are available:

*-sol*

Displays directly a tts-inco SOL file (same filename as the input image).

*-rint [CALfile] [-outdat DATfile]*

Performs radial integration.

If no CALfile is specified, calibration parameters are taken from the image header.

If no DATfile is specified, same name as the input image (but .dat) is used.

*-cal 0/1/2... [dist] [wave] [-outcal [CALfile]]*

Instrumental Parameters Calibration.

The first argument following *-cal* is an integer to select the calibrant substance. It corresponds to: 0= LaB6, 1= Si, 2= first calibrant in config file, 3= second calibrant in cfg file, etc...");

If no dist or wave are specified they are taken from the image header.

Add *-outcal* option to generate a CAL filename with the same name as the input image as long as no CALfile is specified.

*-show*

To open graphical display and do not exit after processing.

(if the first argument is *-help* then the different options are listed)

(as it has been said before, if the only command line argument is a path to an image it will be opened directly)

## Examples:

```
./d2Dplot -macro lab6_180_0003.edf -cal 0 180 0.3187 -outcal lab6_180.cal
```

Running on Unix or Linux

```
[19:25] 10 vava33.d2dplot.D2Dplot_global [CONFIG] ** LOGGING DISABLED **
```

```
Console logging DISABLED
```

```
[19:28] MACRO MODE ON
```

```
[19:28] Reading img file: lab6_180_capillary_0003.edf
```

```
[19:28] CAL option found, performing LaB6 calibration
```

```
[19:28] Using entered distance 180.000
```

```
[19:28] Using entered wavelength 0.3187
```

```
-----
[19:28] REFINEMENT RESULTS:
-----
```

```
CenterX=1023.43250
```

```
CenterY=1023.45001
```

```
S-D_distance=181.57628
```

```
ROT=35.75453
```

```
TILT=-1.25287
-----
```

```
[19:28] Writting output CAL file: lab6_180.cal
```

```
./d2Dplot -macro lab6_180_0003.edf -rint lab6_180.cal
```

Running on Unix or Linux

```
[19:25] 10 vava33.d2dplot.D2Dplot_global [CONFIG] ** LOGGING DISABLED **
```

```
Console logging DISABLED
```

```
[19:26] MACRO MODE ON
```

```
[19:26] Reading img file: lab6_180_0003.edf
```

```
[19:26] RINT option found, performing Radial Integration
```

```
[19:26] Using integration parameters from CAL file: lab6_180.inp
```

```
[19:26]
```

```
[19:26] x-beam center: 1023.430
```

```
[19:26] y-beam center: 1023.450
```

```
[19:26] distance: 181.576
```

```
[19:26] wavelength: 0.3187
```

```
[19:26] tilt rotation: 35.8
```

```
[19:26] angle of tilt: -1.25
```

```
[19:26]
```

```
[19:26] t2ini: 0.000
```

```
[19:26] t2fin: 23.866
```

```
[19:26] stepsize: 0.0236
```

```
[19:26] start azimuth: 0.0
```

```
[19:26] end azimuth: 360.0
```

```
[19:26] subadu: -9.5
```

```
[19:26]
```

```
[19:26] Writting output DAT file: lab6_180_0003.dat
```



## 7. Image formats info

### D2D format

This is a ASCII-header *d2Dplot* format followed by a binary data part using and encoding of unsigned shorts (2-byte little-endian unsigned integers), similar to EDF or IMG formats but with custom header items. It looks like this:

```
{
ByteOrder = LowByteFirst
DataType = UnsignedShort
DataSize = 8388608
Dim_1 = 2048
Dim_2 = 2048
Beam_center_x = 1023.66
Beam_center_y = 1024.22
Pixelsize_x = 79.00
Pixelsize_y = 79.00
Ref_distance = 199.61
Ref_wave = 0.4246
Det_tiltDeg = 0.000
Det_rotDeg = 0.000
Scan_omegaIni = 0.0
Scan_omegaFin = 0.0
Scan_acqTime = -1.0
EXZMargin =0
EXZThreshold =1
EXZdetRadius=1024
EXZpol1 =976 982 957 1013 964 1048 986 1073 1016 1081 1059 1075 1085 1039 1088 998 1057 969 1005 958 313 263 292 282
EXZarc1=1325 1067 52 14
EXZarc2=507 1167 33 8
}
```

### BIN format

This is a pure binary *d2Dplot* format. There are 60 bytes of header followed by the diffraction data (starting at byte 61) using an encoding of signed shorts (2-byte little-endian signed integers). The header info is:

Integer (4 bytes)	dimension X (image "columns") in pixels
Integer (4 bytes)	dimension Y (image "rows") in pixels
Real (4 bytes)	Image scale factor
Real (4 bytes)	Beam X (in pixels)
Real (4 bytes)	Beam Y (in pixels)
Real (4 bytes)	Pixel size X (microns)
Real (4 bytes)	Pixel size Y (microns)
Real (4 bytes)	Sample-to-detector distance (mm)
Real (4 bytes)	Wavelength (Angstrom)
Real (4 bytes)	Omega initial (degrees)
Real (4 bytes)	Omega final (degrees)
Real (4 bytes)	Acquisition time (seconds)

## EDF format

ESRF Data Format. (search the ESRF web page for more info, e.g. <http://www.esrf.eu/computing/scientific/SAXS/doc/SaxsKeywords/SaxsKeywords.pdf>). There are different implementations of the format, the one supported by *d2Dplot* looks like this:

```
{
HeaderID = EH:000001:000000:000000 ;
ByteOrder = LowByteFirst ;
DataType = UnsignedShort ;
Size = 8388608 ;
Dim_1 = 2048 ;
Dim_2 = 2048 ;
beam_center_x = 1023.66 ;
beam_center_y = 1024.22 ;
pixelsize_x = 79.00 ;
pixelsize_y = 79.00 ;
ref_distance = 199.61 ;
ref_wave = 0.4246 ;
scan_type = mar_ct (-1.0,) ;
}
(binary data)
```

## IMG format

ADSC-style IMG files [Arvai, A. J., & Nielsen, C. (1983). ADSC Quantum-210 ADX]. The ones supported looks like this:

```
{
HEADER_BYTES= 512;
TYPE=unsigned_short ;
BYTE_ORDER=little_endian;
SIZE1=2048;
SIZE2=2048;
DISTANCE= 199.610 ;
PIXEL_SIZE= 0.079000 ;
WAVELENGTH=0.424600;
BEAM_CENTER_X=80.87;
BEAM_CENTER_Y=80.91;
}
(binary data)
```

## GFRM format

Bruker, A. X. S. Area Detector Frame Format [e.g. GADDS detector, Bruker, A. X. S. "General Area Detector Diffraction System (GADDS) User Manual." *Madison, WI* 4 (1999)]

## SPR format

"Spreadsheet" format. Table of intensities in ASCII format with the image pixel size (X Y) in the first line.

```
2048 2048
1.78000E+02 1.61000E+02 1.73000E+02 1.86000E+02 2.23000E+02 2.57000E+02 ... (... 2048 columns)
1.23000E+02 2.36000E+02 1.77000E+02 1.56000E+02 1.88000E+02 2.56000E+02 ... (... 2048 columns)
... (2048 rows).
```

## **TIFF format**

TIF image format.

## **CBF format**

DECTRIS Pilatus (Henrich *et al.* 2009) image format.

[[https://www.dectris.com/technical\\_pilatus.html?.../pilatus/](https://www.dectris.com/technical_pilatus.html?.../pilatus/)]

*d2Dplot* only supports the following implementation of CBF:

- Compression: CBF\_BYTE\_OFFSET
- Content-Transfer-Encoding: BINARY
- X-Binary-Element-Type: “signed 32-bit integer”
- X-Binary-Element-Byte-Order: LITTLE\_ENDIAN

## 8. Other file formats info

### Database (DB) format

The database files (\*.DB) contain crystallographic information of compounds. They are plain text files with an entry like this one for each of the compounds:

```
#COMP: Lanthanum hexaboride
#NAMEALT: here alternative names can be introduced (will be used in the name search filter)
#NAMEALT: there can be more than one line like this
#FORMULA: La B6
#CELL_PARAMETERS: 4.1569 4.1569 4.1569 90.000 90.000 90.000
#SPACE_GROUP: P m 3 m
#REF: National Institute of Standards and Technology
#COMMENT: Any comment regarding the entry (temperature, pressure, etc...) can be entered here.
#COMMENT: Also multiple comment fields can be added.
#LIST: H K L dsp Int
  1  0  0  4.15760  13.60
  1  1  0  2.93990  21.83
  1  1  1  2.40040  42.36
  2  0  0  2.07880  56.99
  2  1  0  1.85930  11.83
  2  1  1  1.69730  5.82
  2  2  0  1.46990  0.24
  2  2  1  1.38590 100.00
  3  0  0  1.38590  54.08
  3  1  0  1.31470  67.94
  3  1  1  1.25360  4.49
```

Different compounds are separated by a blank line.

Of all the fields, the only ones that are really required are the compound name (#COMP) and the d-spacing list (#LIST), which can be also introduced without intensities.

Compounds in the database can be added manually with a text editor or by using the database module of *d2Dplot* (add/edit compound).

### Excluded zone (EXZ) format

The excluded zone file itself have comments explaining the three possible fields defining excluded zones. It looks like this:

```
! Excluded zones file for: /home/ovallcorba/lab6_29p2_200_coll_0000.edf
EXZmargin=0
EXZthreshold=0
EXZdetRadius=1024
EXZpol1=997 581 889 385 646 530 847 510 855 592
EXZarc1=1325 1067 52 14
!
! EXZmargin      Margin of the image in pixels (if any)
! EXZthreshold   Pixels with Y<threshold will be excluded
! EXZdetRadius   To exclude corners of the image in case detection area is circular(radius in px)
! EXZpol#        Sequence of pixels (X1 Y1 X2 Y2 X3 Y3...) defining a polygonal shape
! EXZarc#        Arc-shape defined as: ArcCenterX ArcCenterY ArcHalfRadialWthPx ArcHalfAzimWthDeg
```

## 9. References

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## 10. Miscellaneous

### Release notes

*d2Dplot* development started on 2013 as a tool to visualize the orientation search results of microvolumes (*tts*-INCO and related methodologies, [Rius \*et al.\* IUCrJ. 2015; 2, 452-463](#)) and as complement to develop the technique. The program has grown a lot since then, and while it remains basically a tool to visualize diffraction images it may be useful and interesting for a general usage. This is why after a little polishing it has been made available for use.

Feedback to the author would be greatly appreciated. Also, if you find interesting to add a certain functionality ask me and I will try my best.

*d2Dplot* is completely programmed with Java™ ([www.java.com](http://www.java.com)) using jdk version 1.6. (Oracle License: <http://www.oracle.com/technetwork/java/javase/downloads/jdk-6u21-license-159167.txt>).

Major recent changes:

- 1811. Added the possibility to use custom calibrants added in the config file.
- 1805. Redesigned database module for easier editing. Added import CIF and the possibility to calculate the reflections from symmetry.
- 1805. Added Azimuthal plot, fast viewer and SC to INCO.
- 1805. Changed convention of tilt/rot.
- 1805. Added background support to peak integration.
- 1704. Reading of TIFF and CBF formats.
- 1704. Read/write tilt and rot from/to EDF headers.

The following 3<sup>rd</sup> party libraries have been used:

- Commons Math. <https://commons.apache.org/proper/commons-math/>  
Apache License: <http://www.apache.org/licenses/LICENSE-2.0>
- MigLayout. <http://www.miglayout.com>  
BSD license: [http://directory.fsf.org/wiki/License:BSD\\_4Clause](http://directory.fsf.org/wiki/License:BSD_4Clause)
- ImageJ 1.50i. <https://imagej.nih.gov/ij/index.html>  
Public-domain: <https://imagej.net/Licensing>.

(No changes on the source codes of these libraries have been made, you can download the source codes for these libraries at their respective websites).



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Citation of [O. Vallcorba & J. Rius. d2Dplot: 2D X-ray diffraction data processing and analysis for through-the-substrate microdiffraction J. Appl. Cryst. 2019, 52, 478-484](#) would be greatly appreciated when this program helped to your work.

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The purpose of the database system implemented in the program is the creation of a personal compound database by the users. The authors of the program (or their institutions) take no responsibilities in respect of where the data is taken from or its correctness. The default DB is a small selection of 60 compounds coming from different sources. Each entry contains the reference from where it has been taken (with the respective authors). If any of these entries should be removed (for whatever reason) please contact the author and they will be removed immediately.

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